

${\cal N}$ Goldstini

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Motivation

- \rightarrow Supersymmetry is spontaneously broken.
- $\rightarrow~$ In the low energy it is then non-linearly realized.
- $\rightarrow~$ This can be done with constrained superfields.

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N = 1 goldstino

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Nilpotent goldstino superfield

Rocek '78, Lindstrom, Rocek '79, Casalbuoni, De Curtis, Dominici, Feruglio, Gatto '89

We can break SUSY with a chiral superfield

$$X = A + \sqrt{2}\,\theta G + \theta^2 F.$$

- SUSY broken: $\delta G_{\alpha} = -f \epsilon_{\alpha} + \cdots$
- ▶ In the formal limit $m_A \rightarrow \infty$ the scalar decouples.
- This is described by imposing the superspace constraint

$$X^2 = 0
ightarrow X = rac{G^2}{2F} + \sqrt{2}\, heta^lpha G_lpha + heta^lpha heta_lpha F.$$

Volkov–Akulov from X

The simplest supersymmetric Lagrangian is

$$\mathcal{L} = \int d^4 \theta X \overline{X} + \left\{ \int d^2 \theta f X + c.c. \right\}.$$

In component form (after eliminating F) we get

$$\mathcal{L} = -f^2 + i\partial_m \overline{G} \overline{\sigma}^m G + \frac{1}{4f^2} \overline{G}^2 \partial^2 G^2 - \frac{1}{16f^6} G^2 \overline{G}^2 \partial^2 G^2 \partial^2 \overline{G}^2.$$

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Komargodski, Seiberg '09

- Fermion D.O.F. \neq Boson D.O.F.
- $\bullet \ \delta G_{\alpha} = -f\epsilon_{\alpha} (i/2f)\sigma_{\alpha\dot{\alpha}}^{m}\bar{\epsilon}^{\dot{\alpha}}\partial_{m}G^{2} + \cdots$

Constraints on matter multiplets

Removing scalars

Brignole, Feruglio, Zwirner '97, Komargodski, Seiberg '09

• We have another chiral superfield ($\overline{D}_{\dot{\alpha}}Y = 0$)

$$Y = y + \sqrt{2}\theta\chi^{Y} + \theta^{2}F^{Y}.$$

We can have

$$XY = 0.$$

This gives

$$y = \frac{G\chi^{Y}}{F} - \frac{G^2}{2F^2}F^{Y}.$$

There are many constraints in the literature, is there a generic prescription?

Removing any selected component

Dall'Agata, Dudas, FF '16

For a generic superfield

$$\boldsymbol{Q} = \boldsymbol{q} + \theta \chi^{\boldsymbol{Q}} + \cdots$$

We propose the (irreducible) constraint

 $X\overline{X}Q=0.$

This removes only the lowest component

$$q = rac{G\chi^{\mathsf{Q}}}{\sqrt{2}F} + \cdots$$

► To remove more components we combine, for example: $X\overline{X}D_{\alpha}Q = 0$ removes also χ^Q .

From irreducible constraints

- $\rightarrow~$ All known constraints are explained.
- \rightarrow We can build new constrained superfields.
- $\rightarrow~$ This corresponds to formal decoupling of eliminated fields.

Future directions

- \rightarrow Relation to other methods.
- $\rightarrow~$ Applications in supergravity and cosmology.
- \rightarrow What happens for N > 1? (The equivalent of $X^2 = 0$?)

N goldstini

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Broken N = 2 in low energy

Cribiori, Dall'Agata, FF '16

For N = 2 we need to describe two goldstini

$$\delta \mathbf{G}_{\alpha} = \xi_{\alpha} \mathcal{F} + \cdots$$
$$\delta \tilde{\mathbf{G}}_{\alpha} = \tilde{\xi}_{\alpha} \mathcal{F} + \cdots$$

and the aux. field \mathcal{F} which gets the vev.

• We propose the N = 2 chiral superfield $\overline{D}_{\dot{\alpha}} \mathcal{X} = 0 = \tilde{D}_{\dot{\alpha}} \mathcal{X}$ with constraints

$$\mathcal{X} \tilde{D}_{\alpha} D_{\beta} \tilde{D}_{\gamma} \mathcal{X} = 0,$$

 $\mathcal{X} \tilde{D}_{\alpha} D_{\beta} D_{\gamma} \mathcal{X} = 0.$

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We "guessed" them by following E. Kandelakis '86.

• The solution gives in N = 2

$$\begin{split} \mathcal{X} = & \frac{G^2 \tilde{G}^2}{4\mathcal{F}^3} + \theta \frac{G\tilde{G}^2}{\sqrt{2}\mathcal{F}^2} + \tilde{\theta} \frac{\tilde{G}G^2}{\sqrt{2}\mathcal{F}^2} + \theta^2 \frac{\tilde{G}^2}{2\mathcal{F}} + \tilde{\theta}^2 \frac{G^2}{2\mathcal{F}} \\ &+ 2\,\theta \tilde{\theta} \frac{G\tilde{G}}{\mathcal{F}} + \sqrt{2}\,\tilde{\theta}^2 \theta G + \sqrt{2}\,\theta^2 \tilde{\theta} \tilde{G} + \theta^2 \tilde{\theta}^2 \mathcal{F}. \end{split}$$

The simplest Lagrangian is

$$\mathcal{L} = \int d^8\theta \mathcal{X} \overline{\mathcal{X}} + f \int d^4\theta \mathcal{X} + c.c.$$

and in component form we find

$$\mathcal{L} = -f^2 + i\partial_m \overline{G} \,\overline{\sigma}^m G + i\partial_m \overline{\tilde{G}} \,\overline{\sigma}^m \tilde{G} + \cdots$$

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Broken N = 2 as constrained N = 1

• We can express X as N = 1 constrained superfields

$$\mathcal{X} = \mathbf{X} \,\overline{\mathbf{H}}^{\alpha} \overline{\mathbf{H}}_{\alpha} + \sqrt{2} \mathbf{X} \,\overline{\mathbf{H}}^{\alpha} \widetilde{\theta}_{\alpha} + \widetilde{\theta}^{2} \mathbf{X}.$$

• Where (for
$$\overline{D}_{\dot{\alpha}}H_{\dot{\beta}}=0$$
)

$$X^2 = 0 , \ X \overline{X} D_{\beta} H_{\dot{\alpha}} = 0 , \ X \overline{X} D^2 H_{\dot{\alpha}} = 0.$$

• We reduce the N = 2 model to the N = 1 Lagrangian

$$\begin{split} \mathcal{L} &= \int d^4\theta \left(\mathbf{X}\overline{\mathbf{X}} - \left| \partial_m \left(\frac{\mathbf{X}\overline{\mathbf{H}}^2}{2} \right) \right|^2 + i\partial_m (\mathbf{X}\overline{\mathbf{H}}^\alpha) \sigma^m_{\alpha\dot{\alpha}} (\overline{\mathbf{X}}\mathbf{H}^{\dot{\alpha}}) \right) \\ &+ f \left(\int d^2\theta \mathbf{X} + \mathbf{c.c.} \right). \end{split}$$

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Future directions

→ General N? (Cribiori, Dall'Agata, FF '16)

 \rightarrow Couple to N = 2 supergravity.

 \rightarrow What is the UV origin? (FF, Kočí, von Unge '16)

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Thank you!

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